



# Dark Matter Searches with PandaX Experiment

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### **Dark Matter Direct Detection**

- Incoming dark matter from the universe
- Scattering with target atom (nucleus and electrons)
  - Goodman and Witten (1985)
  - Energy deposit in the detector





### PandaX: dual-phase xenon TPC

- Paired scintillation (S1) and ionization (S2) signals
  - Precise energy measurement and 3-D position reconstruction
  - Discrimination of nuclear recoil and electron recoil signals





### **PandaX Detectors**

2009

- Increasing the detector sensitive target volume
- Lowering radioactive background



### PandaX-4T (3.7 tonne)



2020-

### **PandaX-4T Experiment**

- Sensitive volume: 3.7 tonne xenon
- 800m<sup>3</sup> high-purity water shielding tank
- Commissioning started in 2020/11





# **Physics Run**

2020/11 – 2021/04	<b>Commissioning (Run 0)</b> 95 days: ~0.6 tonne-year	
2021/07 – 2021/10	<b>Tritium removal</b> xenon distillation, gas flushing, etc	Physics Rup
2021/11 – 2022/05	<b>Physics run (Run 1)</b> 164 days: ~1.0 tonne-year	
2022/09 – 2023/09	<b>CJPL B2 hall construction</b> xenon recuperation, detector upgrade	CIPL-II B2 Hall
Expect to resu	me by the end of 2023	Detector Upgrade

### **Detector Response Model**



### Detector response to DM signals

Signal type	Response model parameters	
DM-nucleon scattering Nuclear Recoil Signal	<ol> <li>Energy reconstruction</li> <li>Light yield</li> </ol>	
DM-electron scattering Electron Recoil Signal	<ol> <li>Charge yield</li> <li>Re-combination &amp; fluctuation</li> </ol>	



### Various calibration approaches

Туре	Source	Method
Electron recoil	<sup>83m</sup> Kr/ <sup>220</sup> Rn	injecting gas source
Nuclear recoil	<sup>241</sup> Am-Be	external source
Nuclear recoil	D-D neutron	external source

# **Energy Reconstruction**

- Light (S1) + Ionization (S2)
- Energy resolution @41.5keV: 6.8%

$$E = 13.7 \text{eV} \times \left(\frac{\text{S1}}{\text{PDE}} + \frac{\text{S2}_{\text{b}}}{\text{EEE} \times \text{SEG}}\right)$$

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## **Nuclear Recoil Calibration**

### Neutron sources

- Deuteron-deuteron and AmBe neutrons
- Combined fitting to get the parameters
  - Light yield
  - Charge yield
  - Fluctuations



### **Electron Recoil Calibration**

- Inject <sup>220</sup>Rn into the detector, uniformly distributed
- Leakage of electron recoil events below NR median
  - $6/1393 = 0.44 \pm 0.18\%$ , response model agrees with data





### **Background composition**



Component	Nominal (evts)	
<sup>3</sup> T (from fit to data)	532 (32)	
Flat ER <sup>*</sup> (18-30keV side band)	492 (31)	
Rn	347 (190)	
Kr	53 (34)	
Material	40 (5)	
pp neutrino	37 (8)	
Xe-136	31 (6)	
Xe-127	8 (1)	
Neutron	0.9 (0.5)	
Neutron-X	0.2 (0.1)	
Surface	0.5 (0.1)	
Accidental	2.4 (0.5)	
B8	0.6 (0.3)	
Sum	1037 (45)	

- Background level at low energy region is 1/4 of PandaX-II
  - Radon is reduced to 1/6
  - Kr is reduced to 1/20
  - Residual tritium is observed



# Signal ROI

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- Fiducial volume: 2.67 tonne xenon
- Exposure: 0.63 tonne-year
- Signal selection criteria
  - S1: 2 135 PE
  - S2raw: > 80 PE
  - S2 < 20000 PE
- Data: 1058 events observed
  - 6 events below NR median, consistent with expectation of  $9.8\pm0.6$  events



### **DM-nucleon Spin-independent Scattering**

- Sensitivity improved from PandaX-II final analysis by 2.6 times at 40 GeV/c<sup>2</sup>
- Dived into previously unexplored territory!



PRL 127, 261802 (2021), Editors' suggestion

### **Luminance of Dark Matter**

Residual weak EM properties: coupling with photons





tree-level

higher-order loop-level

### **Photon-Mediated Interaction**

- Various nuclear recoil signatures
- Dedicated searches of these EM properties



## **Results from Xenon Recoil Data**

- First experimental constraints on DM charge radius
  - 4 orders of magnitude smaller than neutrino
- Strong constraints on other EM properties
  - up to 3 10 times improvement



#### <u>Limits on the luminance of dark matter from xenon recoil</u> <u>data</u>

A direct search for effective electromagnetic interactions between dark matter and xenon nuclei that produce a recoil of the latter is carried out and the first constraint on charge radius of dark matter is derived.

Xuyang Ning, Abdusalam Abdukerim ... Yubo Zhou

Article 17 May 2023



#### Table 1 | Comparison of electromagnetic properties

	dark matter	neutrino	neutron
Charge radius (fm <sup>2</sup> )	<1.9×10 <sup>-10</sup>	[-2.1,3.3]×10 <sup>-6</sup> *	-0.1155 *
Millicharge (e)	<2.6×10 <sup>-11</sup>	<4 ×10 <sup>-35</sup> *	(-2±8)×10 <sup>-22*</sup>
Magnetic dipole (µ <sub>B</sub> )	<4.8×10 <sup>-10</sup>	<2.8×10 <sup>-11</sup> *	-1×10 <sup>-3*</sup>
Electric dipole (ecm)	<1.2×10 <sup>-23</sup>	<2×10 <sup>-21</sup> ′	<1.8×10 <sup>-26</sup> *
Anapole (cm <sup>2</sup> )	<1.6×10 <sup>-33</sup>	~10 <sup>-34 ‡</sup>	~10 <sup>-28 §</sup>

\* Datas are taken from PDG [32]

† Taken from [31] ‡ Taken from [33]

§ Taken from [34]

Nature Vol. 618, Issue 7963, 47-50 (2023)

### **DM-electron Scattering**

- DM and neutrino may have a connection
  - Behave similarly as a heavy neutrino
  - Mono-energetic recoil energy



Bump-hunting on the electron recoil spectrum

 $\chi e \rightarrow e \nu$ 







## **Cross-check XENON1T Excess**

- Effective field models
  - vector, axial-vector
- Result doesn't support the excess



#### Physics News and COMMENTARY

Potential Dark Matter Signal Gives Way to New Limits

October 13, 2022

Results from two leading dark matter experiments—XENONnT and PandaX-4T—rule out an enigmatic signal detected in 2020 and set new constraints on dark matter particle candidates consisting of light fermions, respectively.

Feature on: E. Aprile *et al.* (XENON Collaboration) Phys. Rev. Lett. **129**, 161805 (2022)

Dan Zhang et al. (PandaX Collaboration) Phys. Rev. Lett. **129**, 161804 (2022)



**PRL** 129, 161804 (2022) **Editors' Suggestion** 

### **Tritium Removal**



- Preliminary estimation of tritium level for Run 1
  - Fitting S1 spectrum, keeping S2 blinded
- Extensive tritium measures planned for next run (Run 2)

Period	Run0 Set 4	Run0 Set 5	Run1
Tritium Counts/day/tonne	3.0 ± 0.3	1.6 ± 0.2	0.4 ± 0.1





### **Reducing Detection Threshold**

- Ionization-only (S2-only): no scintillation signal (S1) requirement
  - ROI S2 [60, 200]PE: threshold down to ~100 eV (from ~1 keV)



### **Ionization-only ROI**

- Key challenge: background components
  - No full picture in previous xenon-based experiments
  - Conservative results only



XENON1T PRL

## **Ionization-only Data**

- First complete understanding of all the main backgrounds
  - Micro-discharging (MD)

Small charge, strong run-condition dependence

- Cathode activity

≻ Large charge, large pulse-shape width

- Blind analysis of 0.55 tonne-year exposure
  - 105 events
  - Best-fit background: 95.8  $\pm$  11.3





### **Constraints on DM-electron Scattering**

- Most stringent constraints are derived
  - DM-electron interaction with heavy mediator,  $2 \times 10^{-41}$  cm<sup>2</sup>
  - Freeze-out and Freeze-in



**PRL** 130, 261001 (2023), **Editors' Suggestion** 

### **Cosmic-ray Boosted Dark Matter**

- Light DM with cosmic ray boosting
- New signature: diurnal modulation due to earth shielding



S.-F. Ge, J. Liu, Q. Yuan, NZ, PRL 126, 091804 (2021)

### **Diurnal Modulation Search**

### PandaX-II data

- Using events below NR median: 25 events (expected 26.6 background)
- Extend the DM search window to sub-GeV
  - Expand to the region beyond the astrophysical and cosmological probes



### **New results from Super-K**

- 20 years' data from Super-Kamiokande PRL 130, 031802 (2023)
- Directional detection of cosmic-ray boosted DM



### Future plan: PandaX-xT

- Next generation liquid xenon experiment
  - with >30 tonne sensitive volume
  - decisive test on WIMP with 200 tonne-year





### Summary



- PandaX-4T is exploring various types of DM
- Novel probes are tested to expand the physics reach
- Run 2 data-taking will start soon
- Planning future PandaX-xT project

# Thank You!

### PandaX



- PandaX: particle and astrophysical xenon detector
  - dark matter, Majorana neutrino, astrophysical neutrino

